

EX18710 (大阪大学推薦課題)

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研究課題名

Possibility for observing space-time distortion via the interaction of ultra-high intense laser pulses with plasmas



1.Introduction



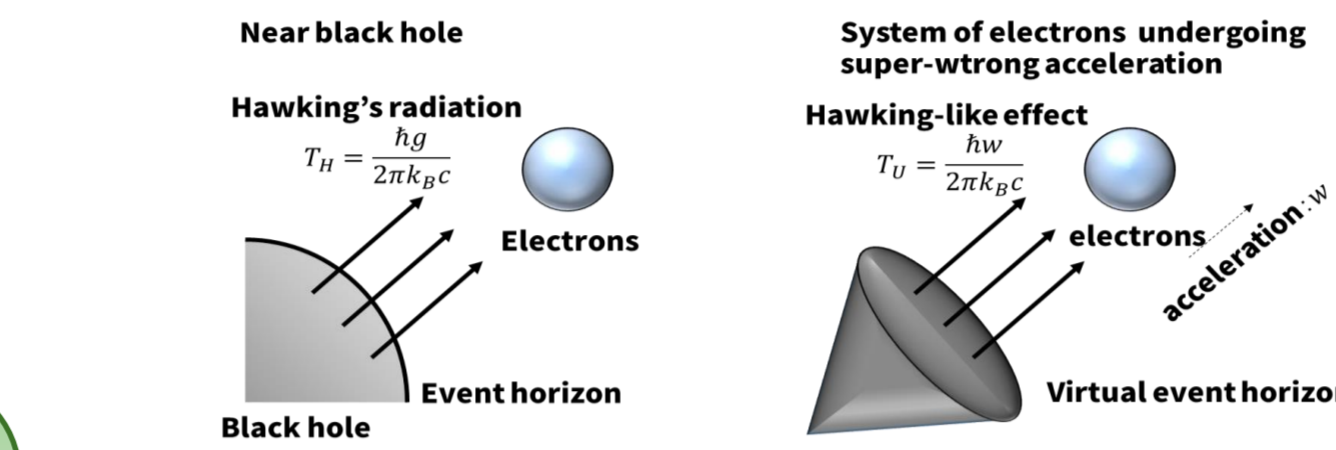
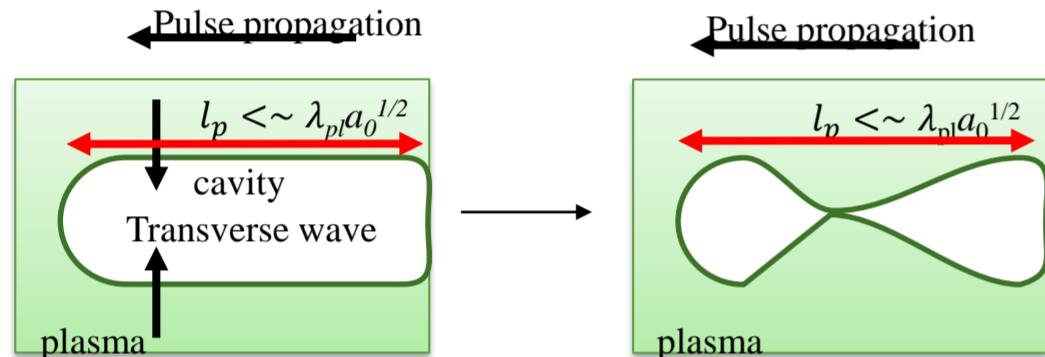
Electrons undergoing super-strong acceleration produced by the laser plasma interactions
Open up new approaches to studying
Space-time physics, vacuum physics,
particle physics, Quantum electrodynamics

For Earth, $g \approx 10^3$ [cm/s²], produces 4×10^{-24} [eV]. In the laser pulse wake, $w > 10^{26}$ g[cm/s²]

For that, we examine the propagation of femtosecond laser pulses with intensities $I=10^{22}$ - 10^{24} W/cm² focused in an underdense plasma by fully relativistic 3D particle-in-cell simulations including ion motion and radiation reaction. The possibility of the detection of Hawking-like effects is discussed.

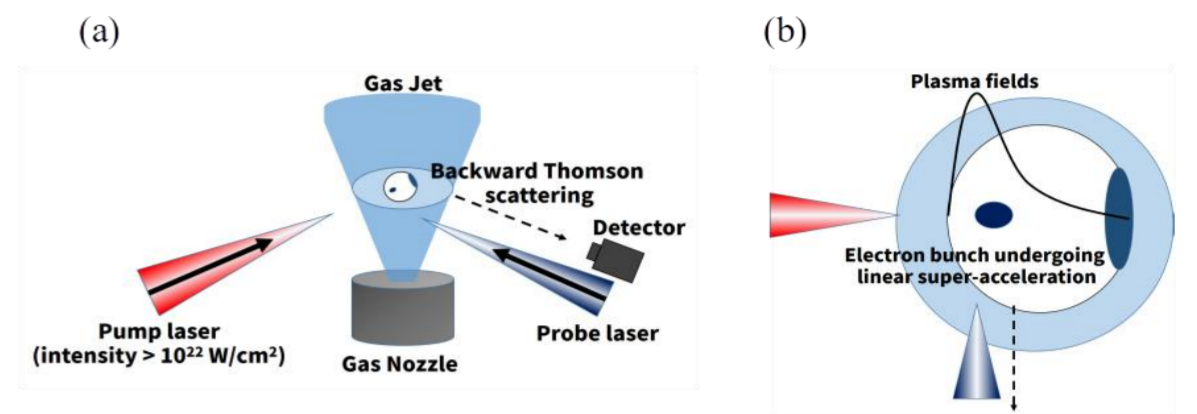
1. Self-focusing.
2. The ion motion.
3. Transverse wave induced by ion motion.
4. Radiation reaction

Transverse plasma wave induced by ion motion



[W. G. Unruh, Phys. Rev. D14, 870 (1976); B. J. B. Crowley et. al. Sci. Rep. 2, 491 (2012).; P. Chen and G. Mourou, Phys. Rev. Lett. 118, 045001 (2017); M. Yano et. al., arXiv:1709.01659 (2017).]

Hawking-like effect: a detector moving with a constant acceleration w sees a boson's thermal bath with its temperature $k_B T_U = \frac{\hbar}{2\pi c} w$ [eV]. For Earth, $g \approx 10^3$ [cm/s²], produces 4×10^{-24} [eV]. In the laser pulse wake, $w > 10^{26}$ g[cm/s²].



Possibility of observing Hawking-like effect by the spectral broadening of Thomson scattering induced by the effect. [M. Yano et. al. (2017)]

The experimental setup using Thomson scattering from the electron bunch undergoing extremely high acceleration is shown. For detection of Hawking-like effect, required strength of the acceleration field is $a_0 \sim 100$.

2.Simulation model

We used Gaussian beam
~30 J- ~3 kJ for 10^{22} - 10^{24} W/cm² and their power of ~3 - ~300 PW
We included ion motion and radiation reaction force

$$E_z(x - x_f, r) = -B_y(x - x_f, r) = \{a_0/d(x)\} \exp[-r^2/d^2(x) - 2x^2/(c\tau)^2 + ikx + i\phi(x - x_f, r)]$$

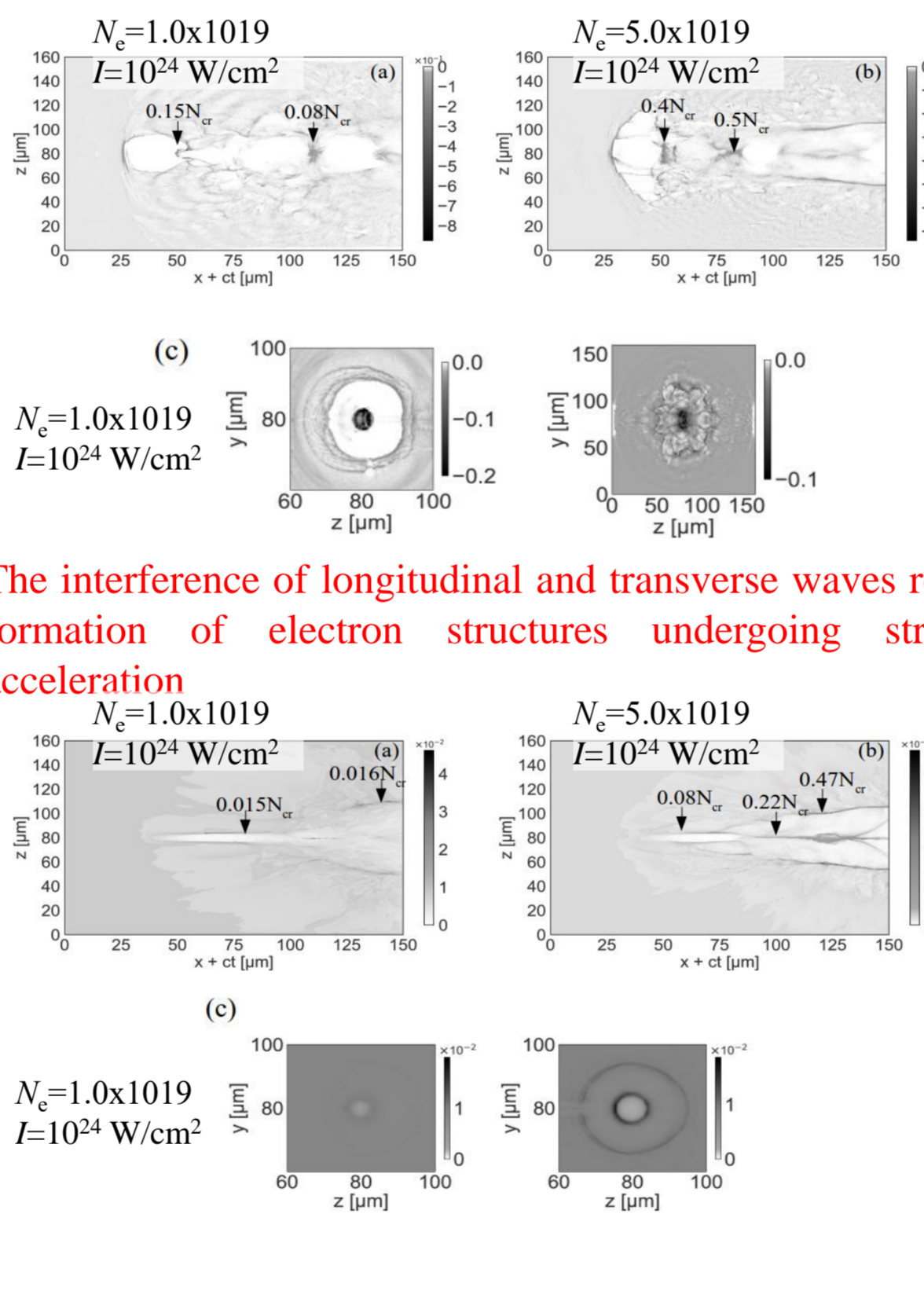
$$\frac{d\vec{p}_i}{dt} = Ze \left(\vec{E} + \frac{\vec{p}_i}{Mc\gamma_i} \times \vec{B} \right)$$

$$F_{RR} \approx -\frac{2e^4}{3m^2c^5} \gamma^2 \vec{v} \left\{ \left(\vec{E} + \frac{1}{c} \vec{v} \times \vec{H} \right)^2 - \frac{1}{c^2} (\vec{v} \cdot \vec{E})^2 \right\}$$

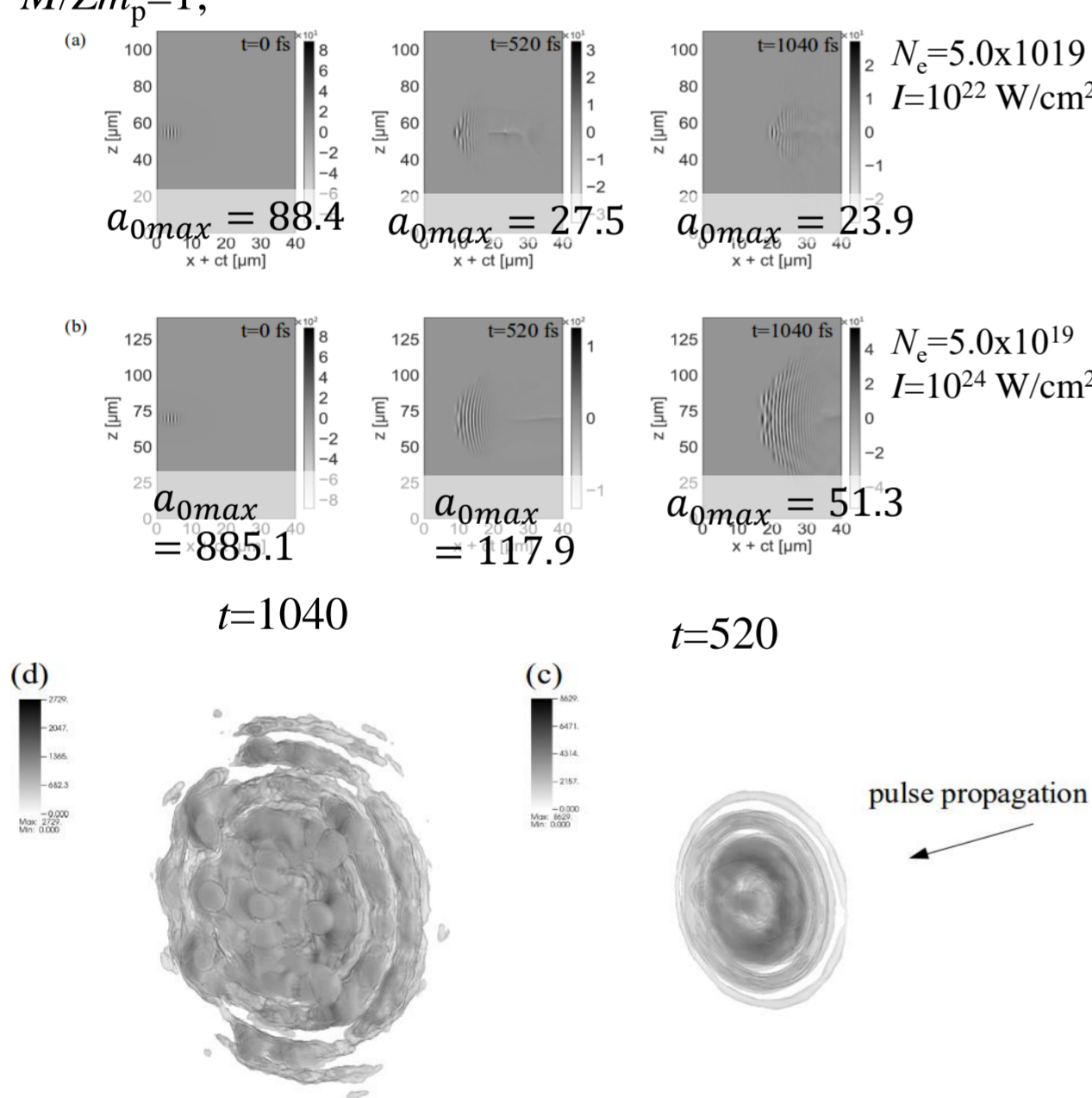
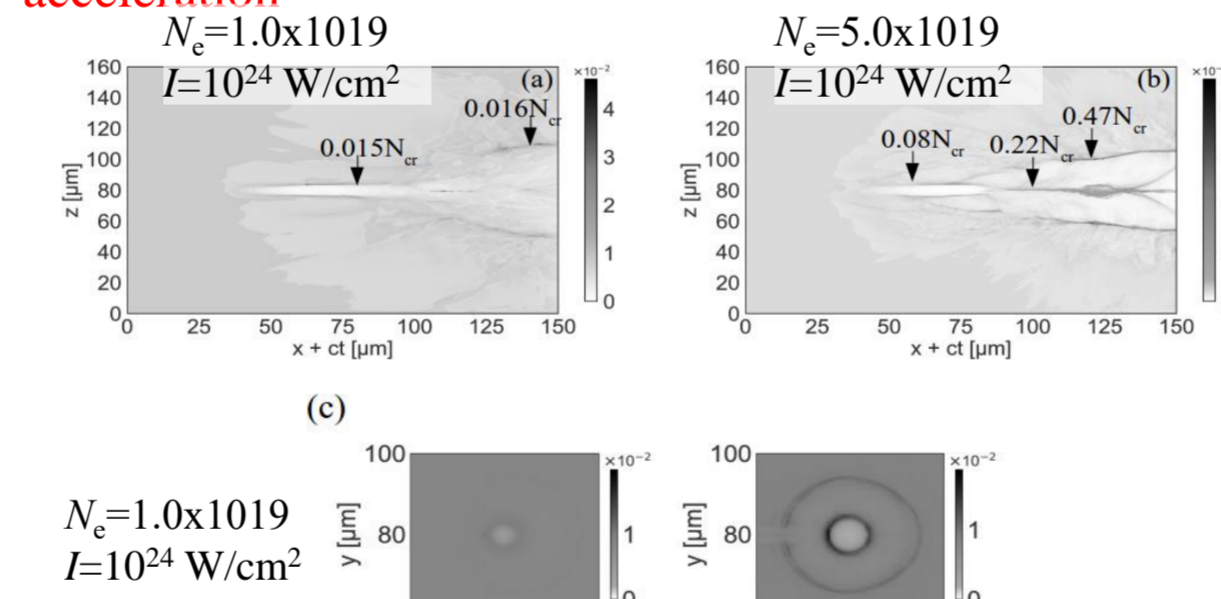
[A. Zhidkov et. al. Phys. Rev. Lett. 88, 185002 (2002). N. Neitz et. al., J. Phys. Conf. Ser. 497, 012015 (2014).;]

3.Simulation results

plasma density $N_e=1.0 \times 10^{19}$ cm⁻³ to $N_e=5.0 \times 10^{20}$ cm⁻³, intensity $I=10^{24}$ W/cm², and mass-to-charge ratio $M/Zm_p=1$;

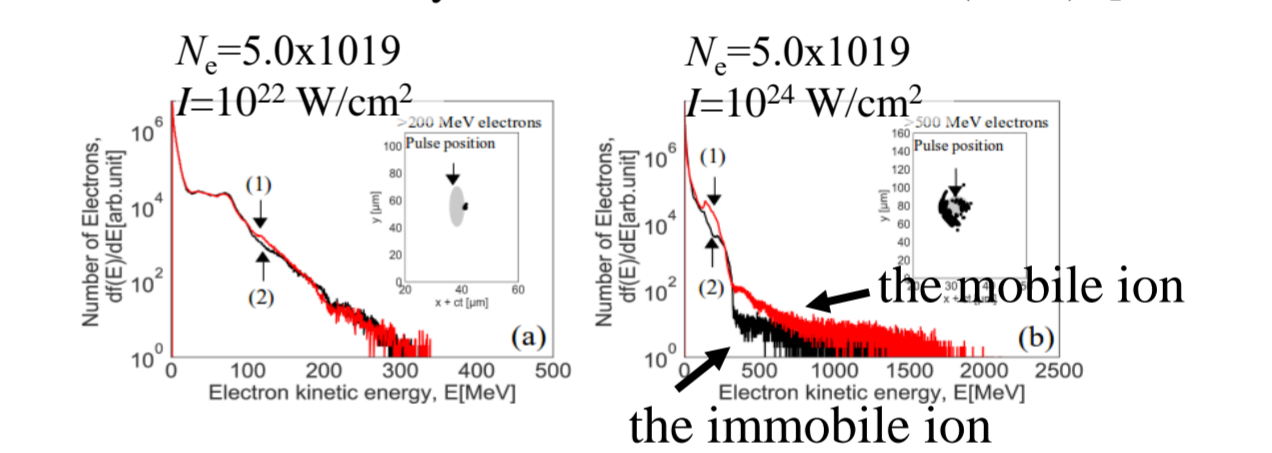


The interference of longitudinal and transverse waves results in the formation of electron structures undergoing strong linear acceleration

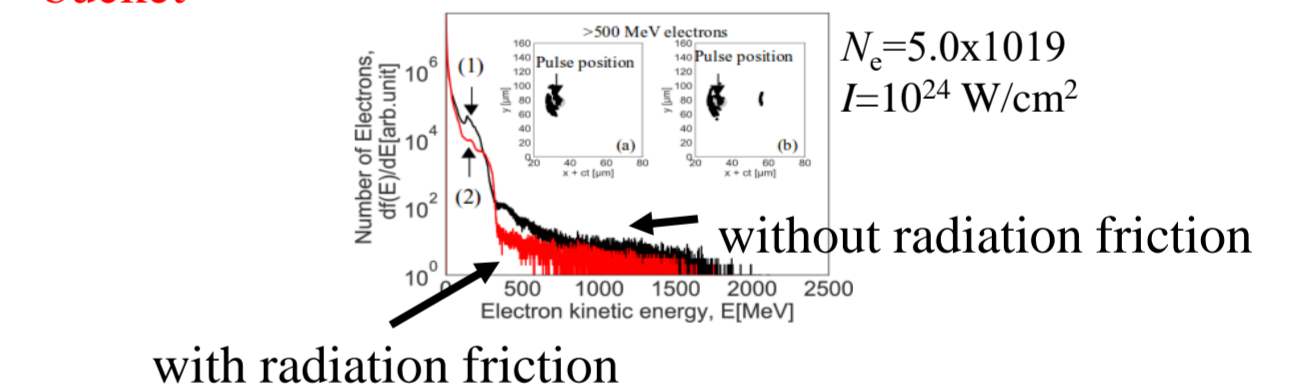


Pulses diffract instantly upon creating vacuum bubble in the plasma

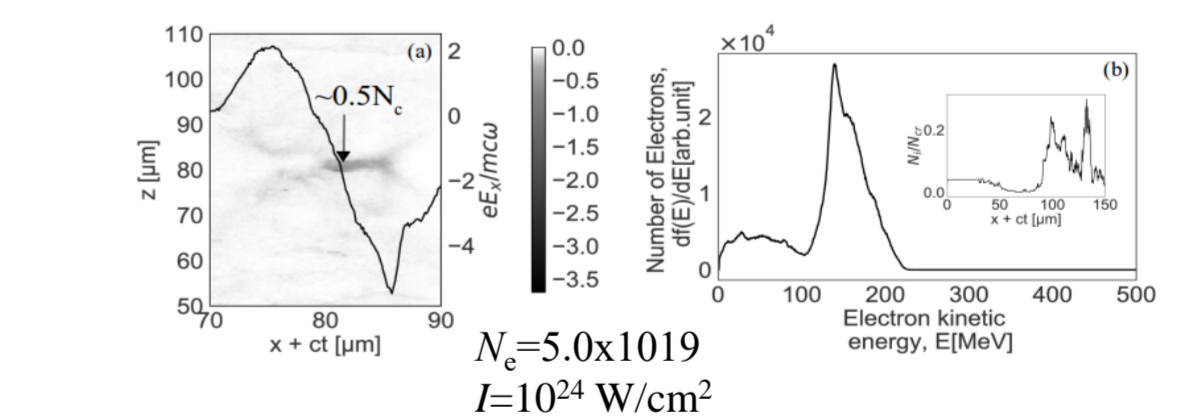
Transverse ion motion originating from a Coulomb explosion results in the formation of a transverse cylindrical plasma wave, which has quite dense electron and ion peaks on the laser axis



The ion motion is not essential to the dynamics of energetic electrons which are mostly located in the first bucket



The radiation reaction dissipates about 10 % of electron energy via Compton scattering of the backward radiation by relativistic electrons



Estimated spectral broadening of the Thomson back scattering light to detect space-time effects gives a value of $\Delta w_s/w_s \sim 0.2$ - 0.5% ($T_U \approx 0.5$ [eV]).

4.Conclusion

- Classical relativistic self-focusing cannot maintain stable propagation of multi-PW class laser pulses with high intensities because pulses diffract instantly upon creating vacuum bubble in the plasma.
- Transverse ion motion originating from a Coulomb explosion results in the formation of a transverse cylindrical plasma wave. The interference of longitudinal and transverse waves results in the formation of electron structures undergoing strong linear acceleration.
- The radiation reaction effects dissipates about 10 % of electron energy via Compton scattering of the backward radiation by relativistic electrons
- Expected spectral broadening of the Thomson back scattering light with an incident angle of 90 degree for a probe pulse from the electrons undergoing super strong acceleration to detect space-time effects gives a value of $\Delta w_s/w_s \sim 0.2$ - 0.5% .

5.Acknowledgement

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